

10.4.2 Quantum Theory of Diamagnetism

In a magnetic field, the generalized momentum \vec{P} of a particle carrying a charge e is

$$P = m \vec{r}' + e \vec{A} = m \vec{v} + e \vec{A} \quad (10.11)$$

where \vec{A} is the *vector potential* defined through Eq. (10.17).

The force \vec{F} acting on the charge e moving with a velocity \vec{v} in an electric and magnetic field is known as *Lorentz force* and is given by

$$\vec{F} = e \left[\vec{E} + (\vec{v} \times \vec{H}) \right] \quad (10.12)$$

where E and H are electric and magnetic field strengths respectively.

The electric and magnetic fields satisfy Maxwell's equations given by

$$\vec{\nabla} \cdot \vec{D} = \rho \quad (10.13)$$

$$\vec{\nabla} \cdot \vec{B} = 0 \quad (10.14)$$

$$\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t} \quad (10.15)$$

$$\vec{\nabla} \times \vec{B} = \mu_0 \vec{J} \quad (10.16)$$

From Eq. (10.14) it follows that \vec{B} can be expressed as the curl of a vector, i.e.

$$\vec{B} = \vec{\nabla} \times \vec{A}, \quad \text{since its divergence is } \nabla \cdot \vec{B} = 0 \quad (10.17)$$

where \vec{A} is the *vector potential of field*.

Thus, Eq. (10.15) takes the form

$$\vec{\nabla} \times \vec{E} = -\frac{\partial}{\partial t} (\vec{\nabla} \times \vec{A})$$

\Rightarrow

$$\vec{\nabla} \times \left(\vec{E} + \frac{\partial \vec{A}}{\partial t} \right) = 0 \quad (10.18)$$

Hence, we may set

$$\vec{E} + \frac{\partial \vec{A}}{\partial t} = -\vec{\nabla} \phi \quad \left[\nabla \times (-\nabla \phi) = 0 \right] \quad (10.19)$$

where ϕ is known as *scalar potential of field*.

We, therefore, get

$$\vec{E} = -\vec{\nabla} \phi - \frac{\partial \vec{A}}{\partial t} \quad (10.20)$$

In terms of vector and scalar potentials, the Lorentz force, given by Eq. (10.12), becomes

$$\vec{F} = m \frac{d\vec{v}}{dt} = e \left[-\vec{\nabla} \phi - \frac{\partial \vec{A}}{\partial t} + (\vec{v} \times \vec{\nabla} \times \vec{A}) \right] \rightarrow \quad (10.21)$$

Now,

$$(\vec{v} \times \vec{\nabla} \times \vec{A})_x = v_y \left(\frac{\partial A_y}{\partial x} - \frac{\partial A_x}{\partial y} \right) - v_z \left(\frac{\partial A_x}{\partial z} - \frac{\partial A_z}{\partial x} \right) \quad (10.22)$$

Since $\frac{\partial}{\partial x} (\vec{v} \cdot \vec{A}) = \frac{\partial}{\partial x} (v_x A_x + v_y A_y + v_z A_z)$ (10.23)

and
$$\frac{dA_x}{dt} = \frac{\partial A_x}{\partial t} + v_x \frac{\partial A_x}{\partial x} + v_y \frac{\partial A_x}{\partial y} + v_z \frac{\partial A_x}{\partial z} \tag{10.24}$$

$$\frac{dA_x}{dt} - \frac{\partial A_x}{\partial t} = v_x \frac{\partial A_x}{\partial x} + v_y \frac{\partial A_x}{\partial y} + v_z \frac{\partial A_x}{\partial z} \tag{10.25}$$

Subtracting Eq. (10.25) from Eq. (10.23), we get

$$\frac{\partial}{\partial x} (\vec{v} \cdot \vec{A}) - \frac{dA_x}{dt} + \frac{\partial A_x}{\partial t} = v_y \left(\frac{\partial A_y}{\partial x} - \frac{\partial A_x}{\partial y} \right) - v_z \left(\frac{\partial A_x}{\partial z} - \frac{\partial A_z}{\partial x} \right) \tag{10.26}$$

Comparing Eq. (10.22) with Eq. (10.26) we find that

$$\left(\vec{v} \times \vec{\nabla} \times \vec{A} \right)_x = \frac{\partial}{\partial x} (\vec{v} \cdot \vec{A}) - \frac{dA_x}{dt} + \frac{\partial A_x}{\partial t} \tag{10.27}$$

Using Eq. (10.27) the components of Lorentz force, given by Eq. (10.21), along the three coordinate axes may be written as

$$\rightarrow F_x = m \frac{dv_x}{dt} = -e \frac{\partial \phi}{\partial x} - e \frac{dA_x}{dt} + e \frac{\partial}{\partial x} (\vec{v} \cdot \vec{A}) \tag{10.28}$$

$$F_y = m \frac{dv_y}{dt} = -e \frac{\partial \phi}{\partial y} - e \frac{dA_y}{dt} + e \frac{\partial}{\partial y} (\vec{v} \cdot \vec{A}) \tag{10.29}$$

$$F_z = m \frac{dv_z}{dt} = -e \frac{\partial \phi}{\partial z} - e \frac{dA_z}{dt} + e \frac{\partial}{\partial z} (\vec{v} \cdot \vec{A}) \tag{10.30}$$

The above three equations may be conveniently put in the following compact vector equation:

$$\frac{d}{dt} (m \vec{v} + e \vec{A}) = \vec{\nabla} [-e\phi + e (\vec{v} \cdot \vec{A})] \tag{10.31}$$

We now put the above equation in the Lagrangian form

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_j} \right) = \frac{\partial L}{\partial q_j} \tag{10.32}$$

where L is the Lagrangian and $q_1 = x, q_2 = y$ and $q_3 = z$.

Clearly, if we assume the Lagrangian to be of the form

$$L = \frac{1}{2} m \vec{v} \cdot \vec{v} - e\phi + e (\vec{A} \cdot \vec{v})$$

and substitute it in Eq. (10.32), we get Eq. (10.31).

The Hamiltonian is given by

$$\mathcal{H} = \vec{p} \cdot \vec{v} - L$$

$$\mathcal{H} = \vec{p} \cdot \left[\frac{1}{m} (\vec{p} - e \vec{A}) \right] - \left[\frac{1}{2} m \left[\frac{1}{m} (\vec{p} - e \vec{A}) \right]^2 + e\phi - \frac{e \vec{A}}{m} \cdot (\vec{p} - e \vec{A}) \right]$$

$$\mathcal{H} = \frac{1}{2m} (\vec{p} - e \vec{A})^2 + e\phi \tag{10.34}$$

$L = T - V$
 $\vec{p} \cdot \vec{v} = p \cdot v \cos \theta = \vec{v} \cdot \vec{A}$
 $\vec{p} \cdot \vec{A} = A \cdot p \cos(-\theta)$
 $\therefore \cos \theta = \cos(-\theta)$

where we have used

$$\vec{v} = \frac{1}{m} (\vec{p} - e\vec{A})$$

from Eq. (10.11).

The stationary state Schrödinger equation describing the motion of a non-relativistic material particle is given by

$$\mathcal{H}\psi = E\psi \quad (10.35)$$

Using Eq. (10.34) and making the substitution $\vec{p} \rightarrow -i\hbar\nabla$, the Schrödinger equation takes the form

$$\begin{aligned} & \frac{1}{2m} (-i\hbar\nabla - e\vec{A})^2 \psi + e\phi\psi = E\psi \\ \Rightarrow & -\frac{\hbar^2}{2m} \nabla^2 \psi + \frac{ie\hbar}{m} [(\nabla \cdot \vec{A})\psi + \vec{A} \cdot \nabla\psi] + \frac{e^2}{2m} A^2 \psi + e\phi\psi = E\psi \end{aligned} \quad (10.36)$$

vector addition

where we have used the vector identity

$$\nabla \cdot (\vec{A}\psi) = (\nabla \cdot \vec{A})\psi + \vec{A} \cdot \nabla\psi$$

From Eq. (10.36) we see that the effect of the magnetic field is to add to the Hamiltonian the terms

$$\mathcal{H}' = \frac{ie\hbar}{m} (\nabla \cdot \vec{A} + \vec{A} \cdot \nabla) + \frac{e^2}{2m} A^2 \quad (10.37)$$

For an electron these terms may be treated as small perturbation. If the magnetic field \vec{B} is uniform, we may choose \vec{A} as

$$\vec{A} = \frac{1}{2} \vec{B} \times \vec{r}$$

\Rightarrow as

$$\vec{A} = \frac{1}{2} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ B_x & B_y & B_z \\ x & y & z \end{vmatrix} \quad (10.38)$$

If the field vector \vec{B} is in the z -direction then we have

$$B_x = B_y = 0; B_z = B$$

and Eq. (10.38) may be written as

$$\vec{A} = \frac{1}{2} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ 0 & 0 & B \\ x & y & z \end{vmatrix}$$

This gives the component of \vec{A} along the three coordinate axes as

$$A_x = -\frac{1}{2}By; \quad A_y = \frac{1}{2}Bx; \quad A_z = 0$$

Bx, By, Bz

Hence,

$$\vec{\nabla} \cdot \vec{A} = \left(\hat{i} \frac{\partial}{\partial x} + \hat{j} \frac{\partial}{\partial y} + \hat{k} \frac{\partial}{\partial z} \right) \cdot (\hat{i}A_x + \hat{j}A_y + \hat{k}A_z)$$

$$\Rightarrow \vec{\nabla} \cdot \vec{A} = \left(\hat{i} \frac{\partial}{\partial x} + \hat{j} \frac{\partial}{\partial y} + \hat{k} \frac{\partial}{\partial z} \right) \cdot \left(-\hat{i} \cdot \frac{1}{2}By + \hat{j} \cdot \frac{1}{2}Bx + 0 \right)$$

$$\Rightarrow \vec{\nabla} \cdot \vec{A} = 0 \quad \left\{ \begin{array}{l} \frac{\partial B_y}{\partial x} = \frac{\partial B_x}{\partial y} = 0 \end{array} \right. \quad (10.39)$$

$\hat{i} \cdot \hat{j} = \hat{j} \cdot \hat{i} = 0$

Again,

$$\vec{A} \cdot \vec{\nabla} = (\hat{i}A_x + \hat{j}A_y + \hat{k}A_z) \cdot \left(\hat{i} \frac{\partial}{\partial x} + \hat{j} \frac{\partial}{\partial y} + \hat{k} \frac{\partial}{\partial z} \right)$$

$$\Rightarrow \vec{A} \cdot \vec{\nabla} = \left(-\hat{i} \cdot \frac{1}{2}By + \hat{j} \cdot \frac{1}{2}Bx + 0 \right) \cdot \left(\hat{i} \frac{\partial}{\partial x} + \hat{j} \frac{\partial}{\partial y} + \hat{k} \frac{\partial}{\partial z} \right)$$

$$\Rightarrow \vec{A} \cdot \vec{\nabla} = -\frac{1}{2}By \frac{\partial}{\partial x} + \frac{1}{2}Bx \frac{\partial}{\partial y}$$

$$\Rightarrow \vec{A} \cdot \vec{\nabla} = \frac{B}{2} \left(x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x} \right) \quad (10.40)$$

and

$$A^2 = A_x^2 + A_y^2 + A_z^2 = \frac{1}{4}B^2y^2 + \frac{1}{4}B^2x^2$$

$$\Rightarrow A^2 = \frac{1}{4}B^2(x^2 + y^2) \quad (10.41)$$

Using Eqs. (10.39) to (10.41), Eq. (10.37) takes the form

$$\mathcal{H}' = \frac{ie\hbar}{m} \cdot \frac{B}{2} \left(x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x} \right) + \frac{e^2}{2m} \cdot \frac{1}{4}B^2(x^2 + y^2) \quad (10.42)$$

The first term on the right-hand side of Eq. (10.42) is proportional to the orbital angular momentum component L_z . In mono-nuclear systems, this term gives rise to paramagnetism. The second term gives for a spherically symmetric system a contribution

$$E' = \frac{e^2 B^2}{12m} \cdot \overline{r^2}, \text{ as } x^2 + y^2 = \frac{2}{3}r^2$$

by first order perturbation energy. The associated magnetic moment is diamagnetic:

$$\mu = -\frac{\partial E'}{\partial B} = -\frac{e^2 B}{6m} \cdot \overline{r^2}$$

Hence, the atomic susceptibility χ_{At} is given by

$$\chi_{At} = \frac{\mu}{H} = -\left(\frac{\mu_0 e^2}{6m} \right) \overline{r^2}, \text{ as } B = \mu_0 H \quad (10.43)$$

and is in agreement with the classical result.

10.4.3 Discussion

On the basis of quantum orbital theory, the mean value of r^2 for an orbit about an effective nuclear charge Z is given by

$$\bar{r}^2 = a_0^2 \cdot \frac{n^2}{Z^2} \left(\frac{5}{2}n^2 - \frac{3}{2}k^2 \right) \quad (10.44)$$

where $a_0 (= 0.528 \times 10^{-10} \text{ m})$ is the radius of the innermost orbit in hydrogen atom, n is the radial quantum number and k , the azimuthal quantum number.

This gives for the contribution of the orbit to the gram atomic diamagnetic susceptibility

$$\chi_A = -2.83 \times 10^{10} \bar{r}^2$$

$$\Rightarrow \chi_A = -0.785 \times 10^{-6} \left[\frac{n^2}{Z^2} \left(\frac{5}{2}n^2 - \frac{3}{2}k^2 \right) \right] \quad (10.45)$$

If the effective nuclear charge is calculated from this expression for helium with two electrons from its observed susceptibility ($\chi_A = -1.9 \times 10^{-6}$), it comes out to be 0.93 which is too small.

To make the expression for susceptibility more accurate, van Vleck and Pauling modified the above expression as

$$\chi_A = -0.785 \times 10^{-6} \left[\frac{n^2}{Z^2} \left(\frac{5}{2}n^2 - \frac{3l(l+1) - 1}{2} \right) \right] \quad (10.46)$$

where l , the orbital quantum number, is equal to $(k - 1)$.

Pauling's calculations necessarily involve a number of approximations. For spherically symmetric atoms, Hartree has devised a method with which the charge distribution satisfying the Schrödinger equation may be worked out more precisely. He has given tables and curves for a number of ions and atoms showing the charge per unit radial distance in a spherical shell of unit thickness.

If $\frac{dN}{dr}$ be the charge in electron unit per unit radial distance, then the number of electrons in the ions is equal to $\int_0^{\infty} \left(\frac{dN}{dr} \right) dr$.

For the diamagnetic susceptibility, he obtained

$$\chi_A = -2.83 \times 10^{10} \int_0^{\infty} \left(\frac{dN}{dr} \right) dr \quad (10.47)$$

This integral can be evaluated graphically (Figure 10.2). Using atomic units for distances, the susceptibility is given by

$$\chi_A = -0.785 \times 10^{-6} \times \text{area under the shaded portion of the curve}$$

For helium this gives

$$\chi_A = -1.89 \times 10^{-6}$$

a value which agrees remarkably closely with the observed one (-1.9×10^{-6}).

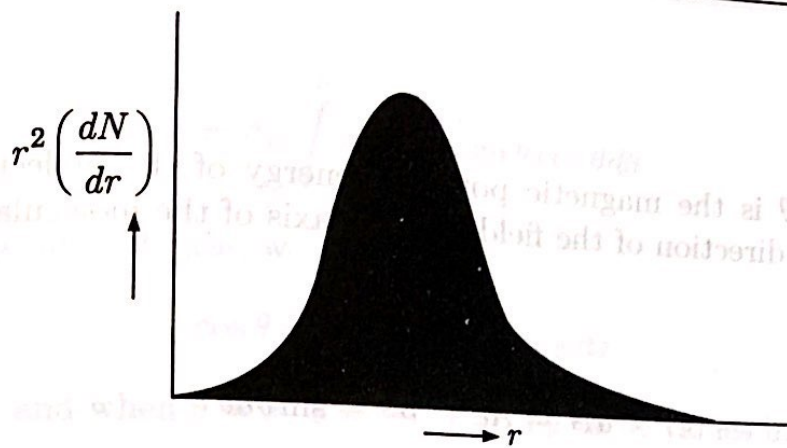


Figure 10.2: Variation of $r^2(dN/dr)$ with r .